Reference energies for valence ionizations and satellite transitions

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Experimental spectrum of water



Ning et al., Chem. Phys. 343 (2008) 19-30

Experimental spectrum of water





Journées "Théories, Modélisation et Simulation", Strasbourg (France), 2023



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The Quest Project



Veril et al., WIREs Comput. Mol. Sci., 11 (2021) e1517

$$\Psi_{0}^{\mathsf{CI}}\rangle = |\Phi_{0}\rangle + \sum_{ia} c_{i}^{a} |\Phi_{i}^{a}\rangle + \sum_{ijab} c_{ij}^{ab} |\Phi_{ij}^{ab}\rangle + \sum_{ijkabc} c_{ijk}^{abc} |\Phi_{ijk}^{abc}\rangle + \dots$$
(1)

The Configuration Interaction Wavefunction

$$|\Psi_{0}^{\mathsf{CI}}\rangle = |\Phi_{0}\rangle + \sum_{ia} c_{i}^{a} |\Phi_{i}^{a}\rangle + \sum_{ijab} c_{ij}^{ab} |\Phi_{ij}^{ab}\rangle + \sum_{ijkabc} c_{ijk}^{abc} |\Phi_{ijk}^{abc}\rangle + \dots$$
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$$\left|\Psi_{0}^{\mathsf{CI}}\right\rangle = \underbrace{\left|\Phi_{0}\right\rangle + \sum_{ia} c_{i}^{a} \left|\Phi_{i}^{a}\right\rangle}_{\mathsf{CIS}} + \sum_{ijab} c_{ij}^{ab} \left|\Phi_{ij}^{ab}\right\rangle + \sum_{ijkabc} c_{ijk}^{abc} \left|\Phi_{ijk}^{abc}\right\rangle + \dots$$
(1)

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(1)

$$\left|\Psi_{0}^{\mathsf{CI}}\right\rangle = \underbrace{\left|\Phi_{0}\right\rangle + \sum_{ia} c_{i}^{a} \left|\Phi_{i}^{a}\right\rangle + \sum_{ijab} c_{ij}^{ab} \left|\Phi_{ij}^{ab}\right\rangle + \sum_{ijkabc} c_{ijk}^{abc} \left|\Phi_{ijk}^{abc}\right\rangle + \dots}_{\mathsf{CISDT}}$$
(1)

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(1)

The selected CI wavefunction

Idea: Select only the most important determinant in each excitation class!

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Garniron et al., J. Chem. Theory Comput. 15 (2019) 3591











Charged excitation energies with selected CI

First ionization potential

$$E_{1 \text{st IP}} = (E_0^N - E_0^{N-1}) \tag{2}$$



Charged excitation energies with selected CI

Valence ionization and satellites

$$E_{\rm IP/Sat} = E_{\rm 1st \ IP} + E_i^{N-1} \tag{2}$$



Charged excitation energies with selected CI

Valence ionization and satellites

$$E_{\rm IP/Sat} = E_{\rm 1st \ IP} + E_i^{N-1} \tag{2}$$



Coupled-cluster hierarchy

Scaling of the coupled-cluster hierarchy

MethodsCCSDCCSDTQScaling N^6 N^8 N^{10}

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$$N^6$$
 N^8 N^{10}

Approximate coupled-cluster hierarchy

 $\begin{array}{c|c} \mbox{Methods} & \mbox{CC2} & \mbox{CC3} & \mbox{CC4} \\ \mbox{Scaling} & \mbox{N^5} & \mbox{N^7} & \mbox{N^9} \\ \end{array}$

The GW approximation (Hedin, Phys. Rev. 139 (1965) A796)



electron removal

The GW approximation (Hedin, Phys. Rev. 139 (1965) A796)



electron removal

The GW approximation (Hedin, Phys. Rev. 139 (1965) A796)



electron removal

Our set of molecules

8 electrons	10 electrons	
BH ₃	Ne, HF, H ₂ O, NH ₃ , CH ₄	
12 electrons	14 electrons	16 electrons
LiF, BeO, C ₂	BF, CO, N_2	CH_2O
18 electrons	20 electrons	22 electrons
Ar, HCl, H ₂ S, F ₂	LiCl	CO ₂ , CS

20 molecules, 49 IPs, 35 Satellites

Statistics for IP







Post-*GW* **self-energies?**

- Able to describe satellites
- Do no deteriorate results for IPs!



Questions?

$$\begin{bmatrix} \mathbf{F} + \mathbf{\Sigma}^{GW} \left(\omega = \epsilon_p^{GW} \right) \right] \varphi_p^{GW} = \epsilon_p^{GW} \varphi_p^{GW}$$

$$\mathbf{\Sigma}^{GW} \left(\omega \right) = \mathbf{W}^{2h1p} \left(\omega \mathbf{1} - \mathbf{C}^{2h1p} \right)^{-1} (\mathbf{W}^{2h1p})^{\dagger}$$

$$+ \mathbf{W}^{2p1h} \left(\omega \mathbf{1} - \mathbf{C}^{2p1h} \right)^{-1} (\mathbf{W}^{2p1h})^{\dagger}$$

$$\lim_{\substack{k = 1 \text{ p conf.} \\ 2h1p \text{ conf.} \\ 2p1h \text{ conf.} \\ } \left\{ \begin{array}{c} \mathbf{F} & \mathbf{W}^{2h1p} \mathbf{W}^{2p1h} \\ \mathbf{W}^{2p1h} & \mathbf{0} \\ \mathbf{W}^{2p1h} & \mathbf{0} \\ \mathbf{W}^{2p1h} \end{array} \right\} \text{ internal space } \mathbf{P}$$

$$\operatorname{external space} \mathbf{Q}$$

Bintrim & Berkelbach, JCP 154 (2021) 041101; Monino & Loos JCP 156 (2022) 231101; Tolle & Chan, JCP 158 (2023) 124123

Hedin's Pentagon



The wonderful equations of Hedin

$$\underbrace{G(12)}_{\text{ten's function}} = G_0(12) + \int G_0(13)\Sigma(34)G(42)d(34)$$

$$\underbrace{\Gamma(123)}_{\text{vertex}} = \delta(12)\delta(13) + \int \frac{\delta\Sigma(12)}{\delta G(45)}G(46)G(75)\Gamma(673)d(4)$$

$$\underbrace{P(12)}_{\text{polarizability}} = -i\int G(13)\Gamma(342)G(41)d(34)$$

$$\underbrace{W(12)}_{\text{screening}} = v(12) + \int v(13)P(34)W(42)d(34)$$

$$\underbrace{\Sigma(12)}_{\text{self-energy}} = i\int G(14)W(13)\Gamma(423)d(34)$$

C 1 1 1

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Gre

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Hedin's Square



The GW approximation

$$\underbrace{G(12)}_{Green's function} = G_0(12) + \int G_0(13)\Sigma(34)G(42)d(34)$$

$$\underbrace{\Gamma(123)}_{vertex} = \delta(12)\delta(13) + \frac{\int \delta\Sigma(12)}{\delta G(45)}G(46)G(75)\Gamma(673)d(4567) + \frac{\int \delta\Sigma(12)}{\delta G(45)}G(46)G(75)\Gamma(673)d(4567) + \frac{1}{2}\int G(12)\Gamma(342)G(21)d(34) = -iG(12)G(21) + \frac{1}{2}\int G(12)\Gamma(342)G(21)d(34) + \frac{1}{2}\int G(12)W(12)\Gamma(423)d(34) = iG(12)W(12)$$
self-energy